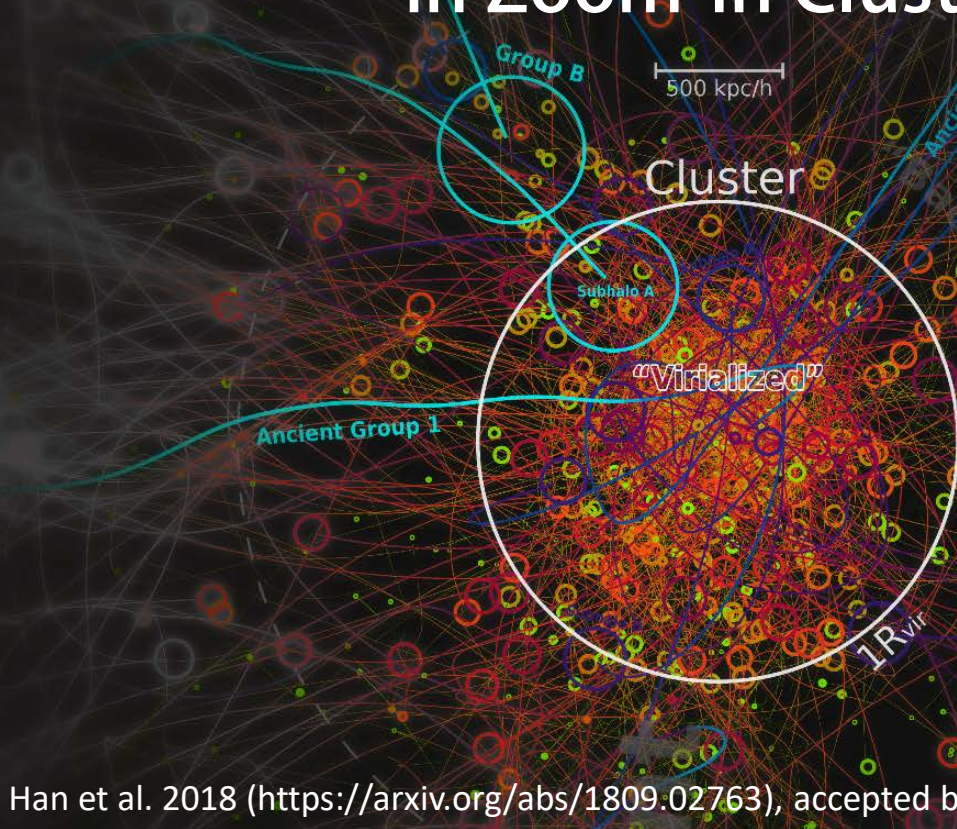


Preprocessing and Mass Evolution of Dark halos in Zoom-in Cluster Simulation

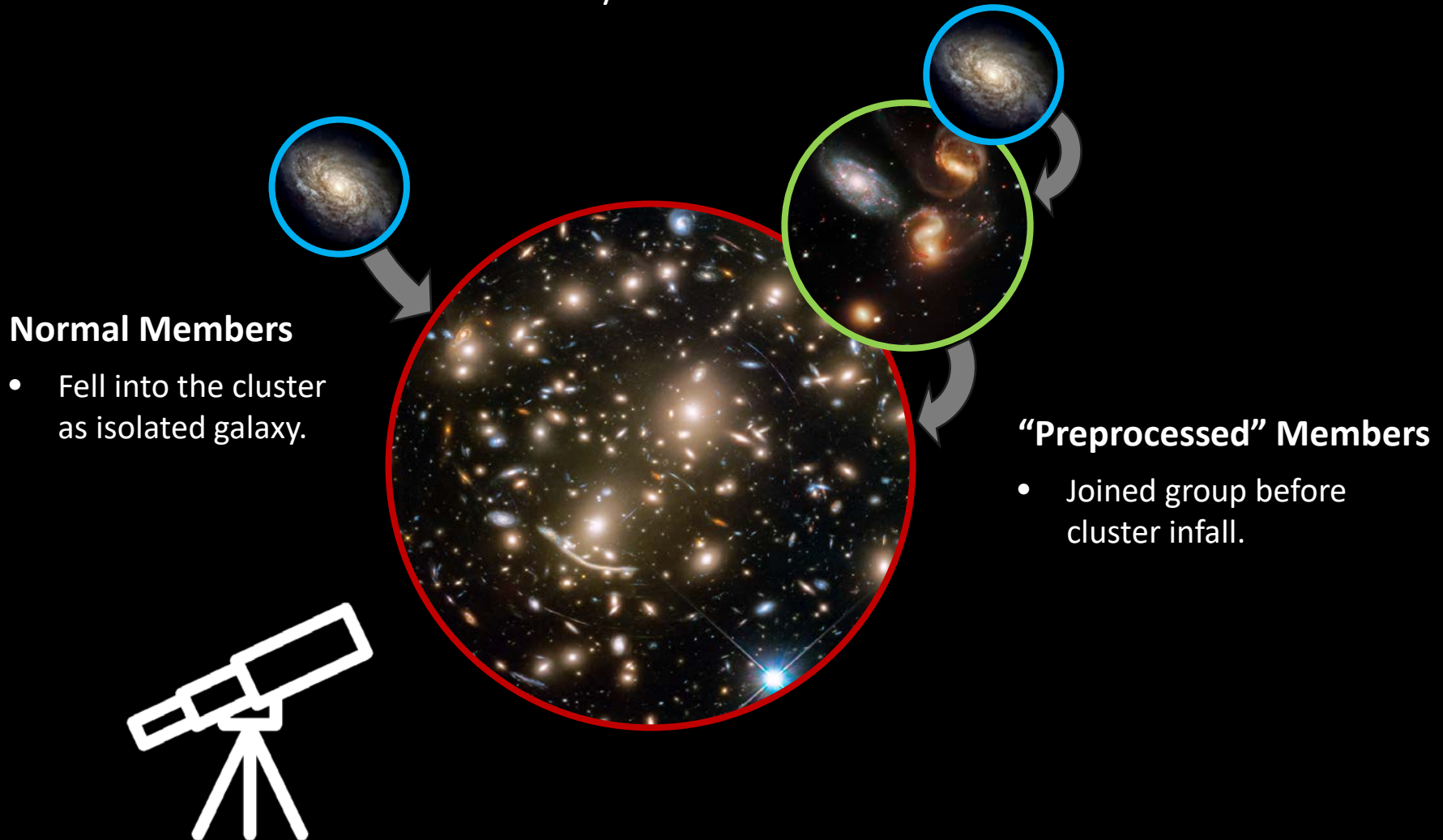


San Han
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Luca Cortese (ICRAR), Barbara Catinella (ICRAR),
Emanuele Contini (Yonsei), Sukyoung Yi (Yonsei)

Concept of Preprocessing

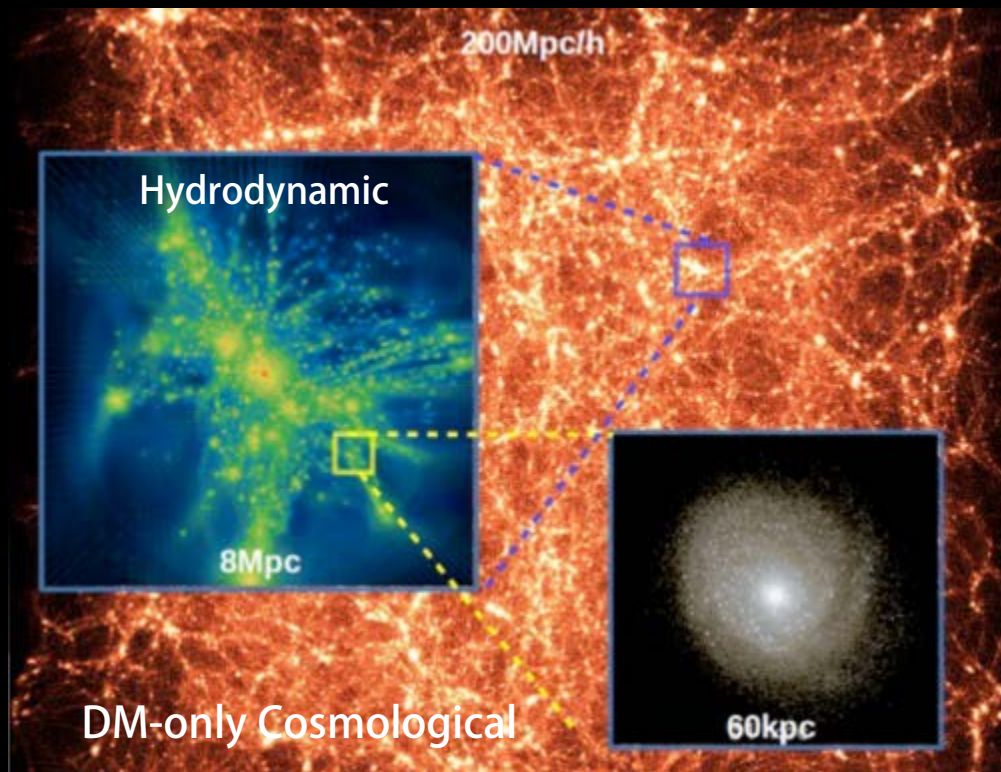
=Environmental effects of smaller systems



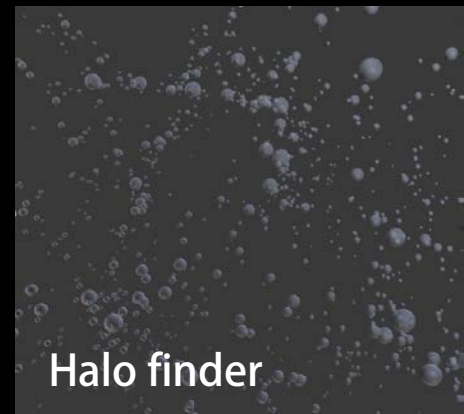
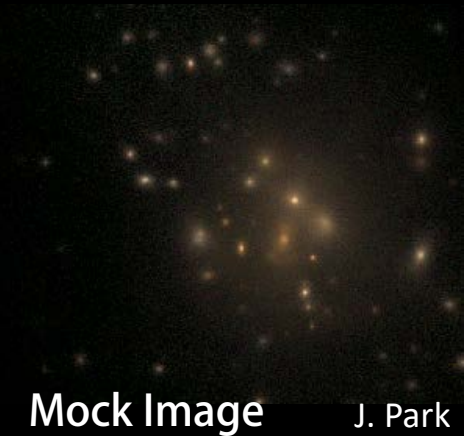
Yonsei Zoom-in Cluster Simulation

Choi & Yi 2017

- 16 regions zoomed into clusters with masses in a range $10^{13.7} - 10^{15} M_{\text{sun}}$.
- Minimum force resolution **0.76kpc**



H. Choi

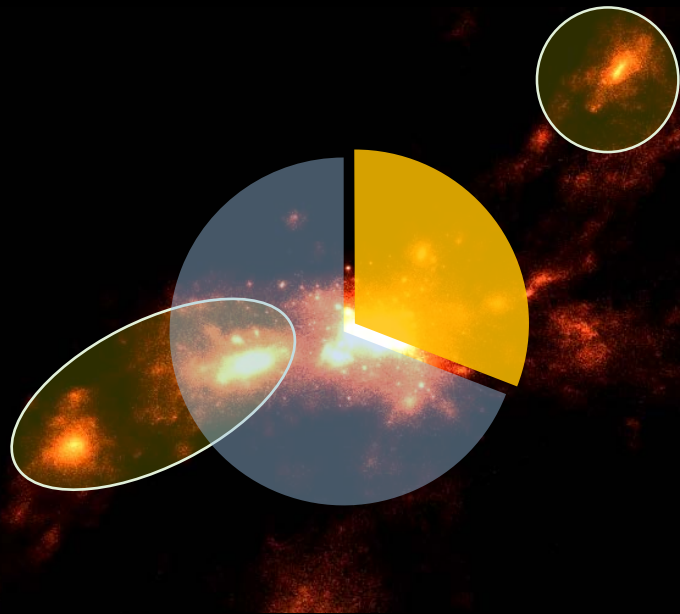


Yonsei Zoom-in Cluster Simulation

Choi & Yi 2017

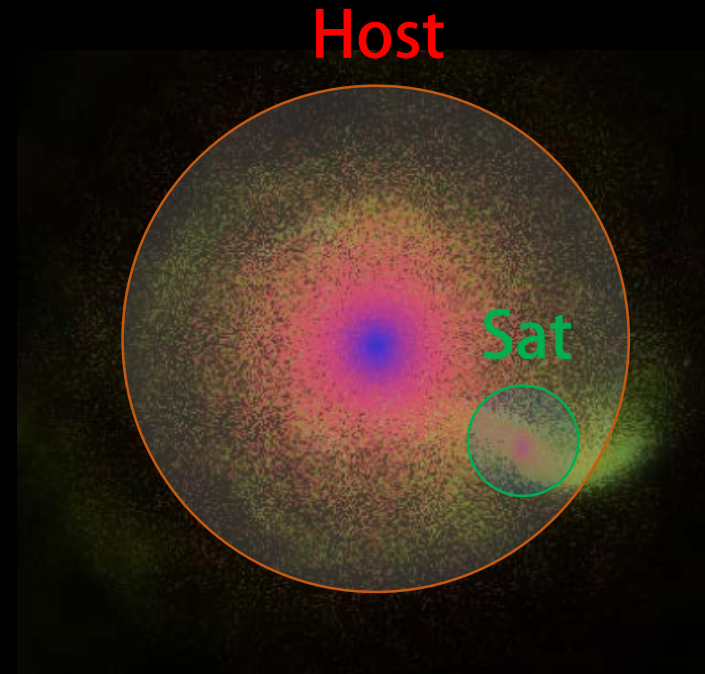


What did we want to know?



Statistics

- What fraction of halos in the cluster have been “preprocessed?”



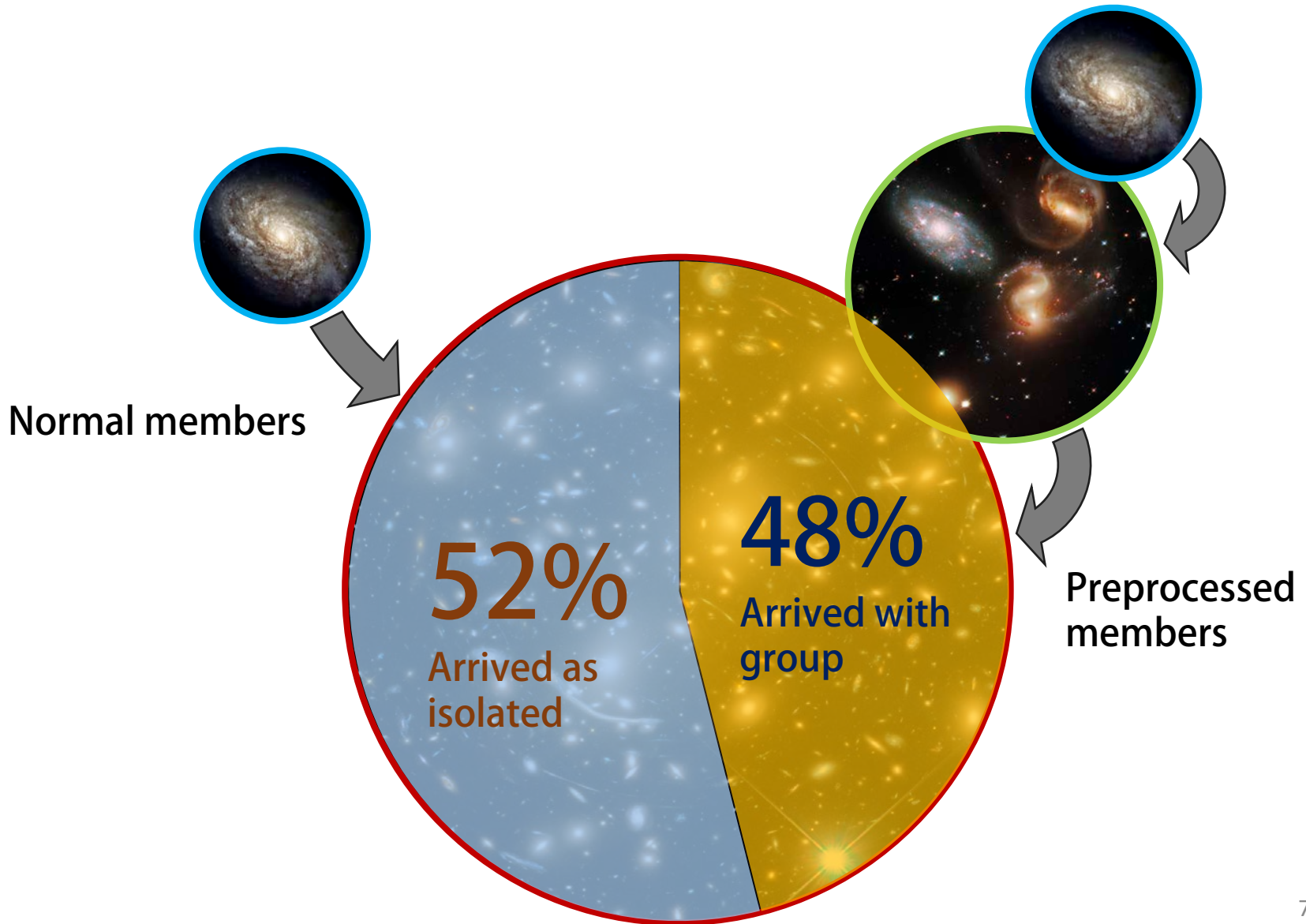
Significance

- How important is the preprocessing?
- When is the effect maximized?

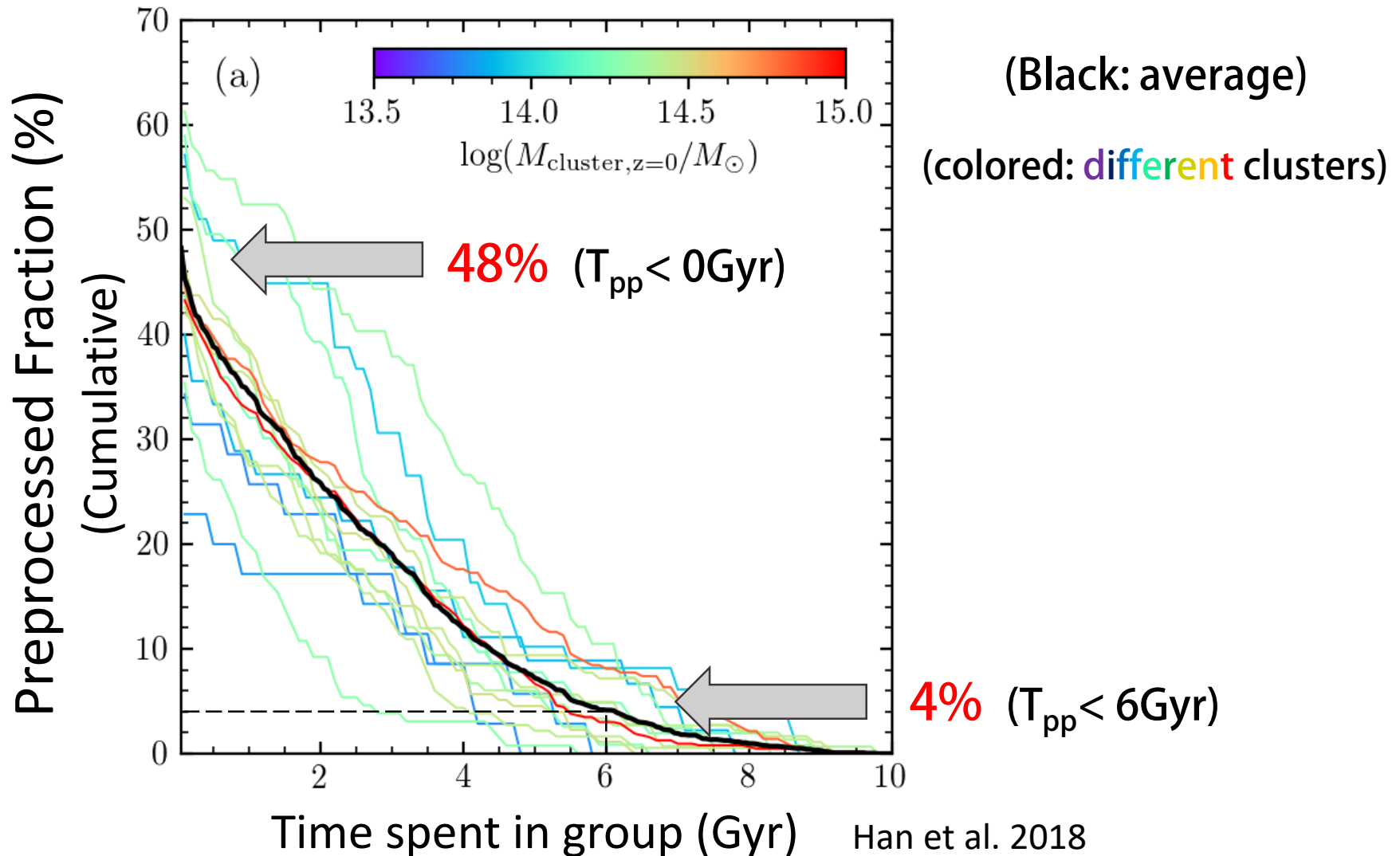
Galaxies vs DM halos

Results

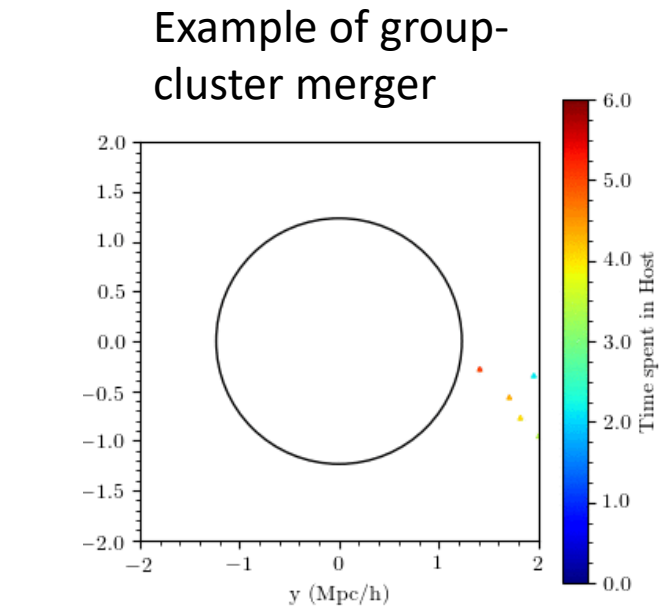
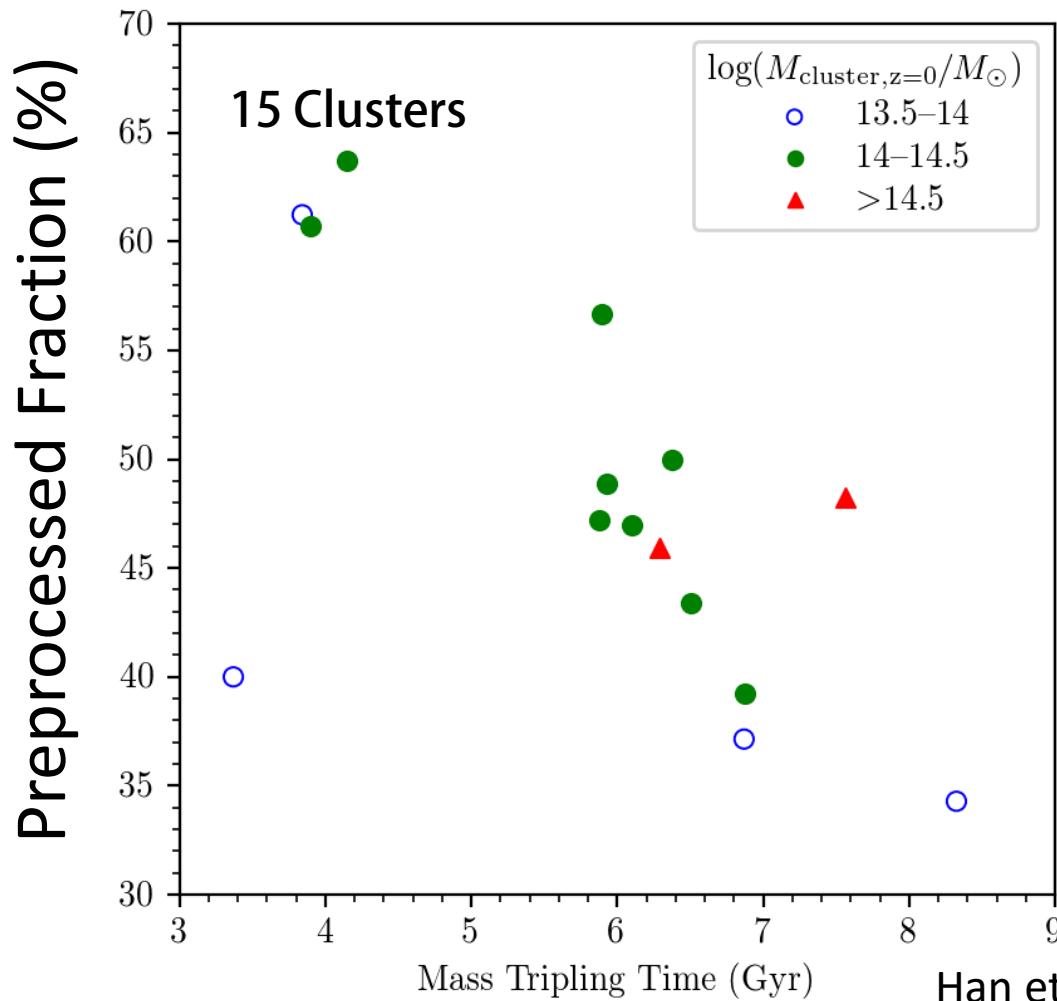
Fraction of preprocessed halos in clusters



Fraction of preprocessed halos in clusters



Recent growth history of the cluster

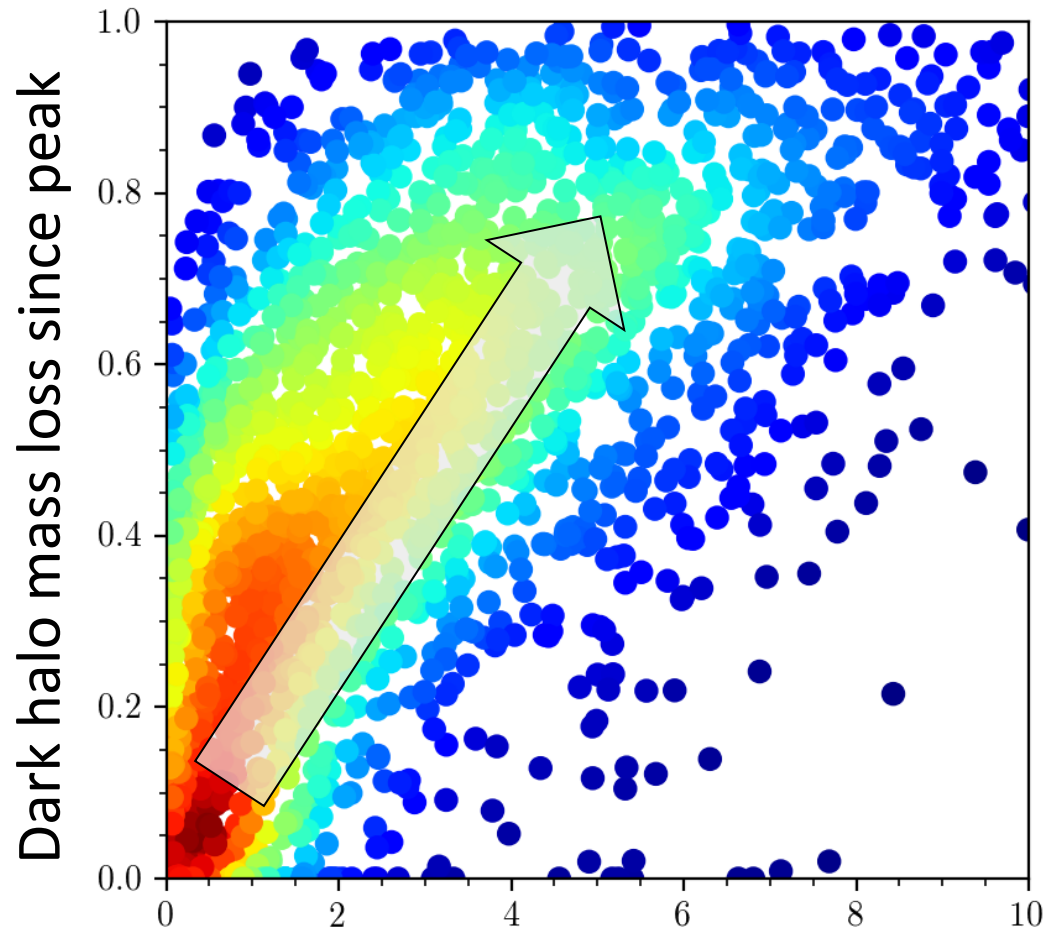


Rapid growth
= Recent group merger



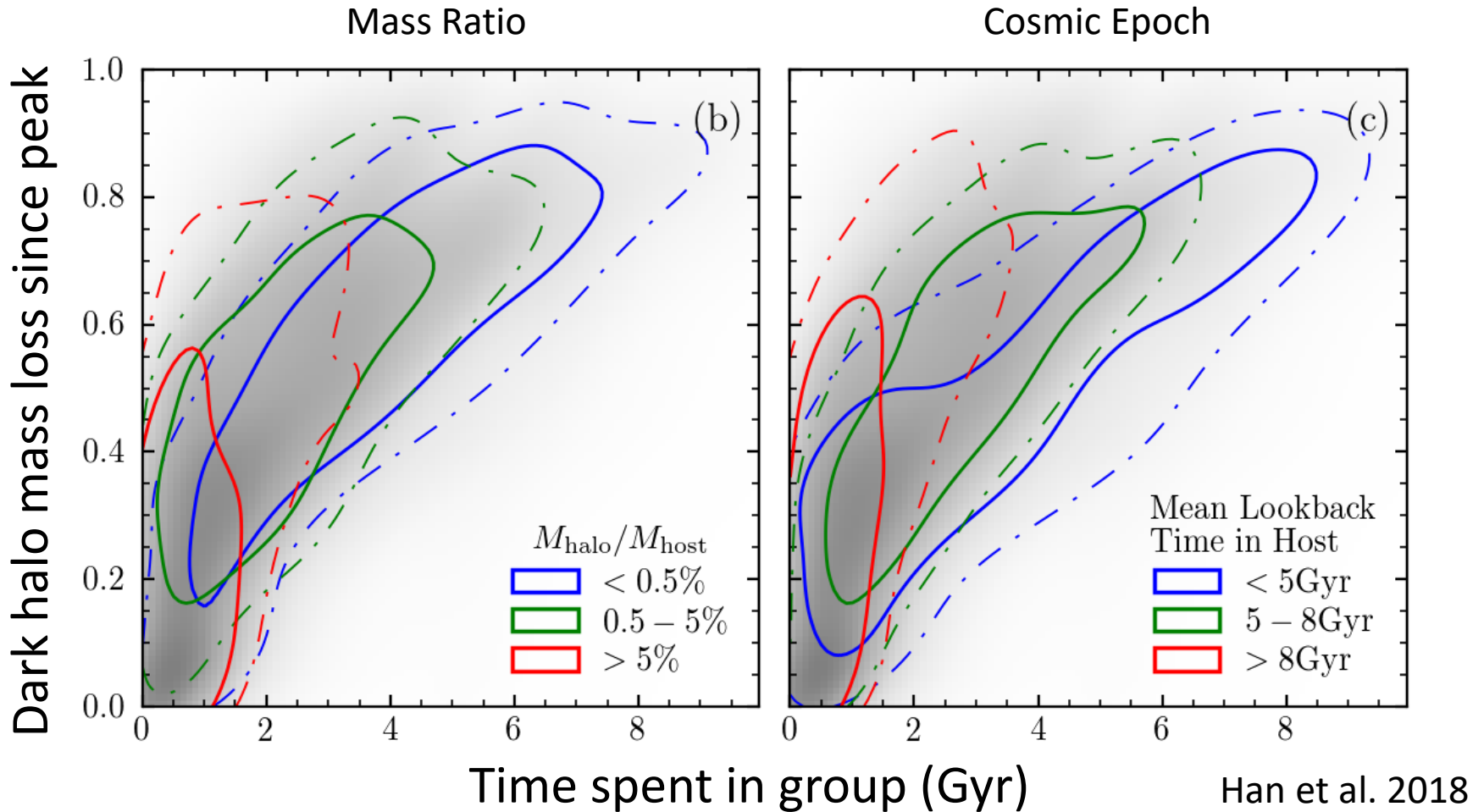
Tidal mass loss of satellites in preprocessing

Color: density of points



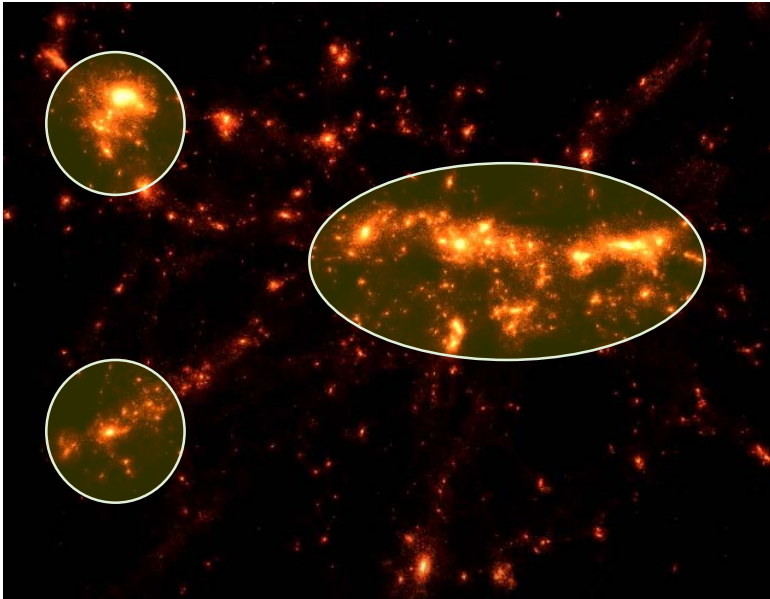
Time spent in group (Gyr) Han et al. 2018

Variation of tidal mass loss rate



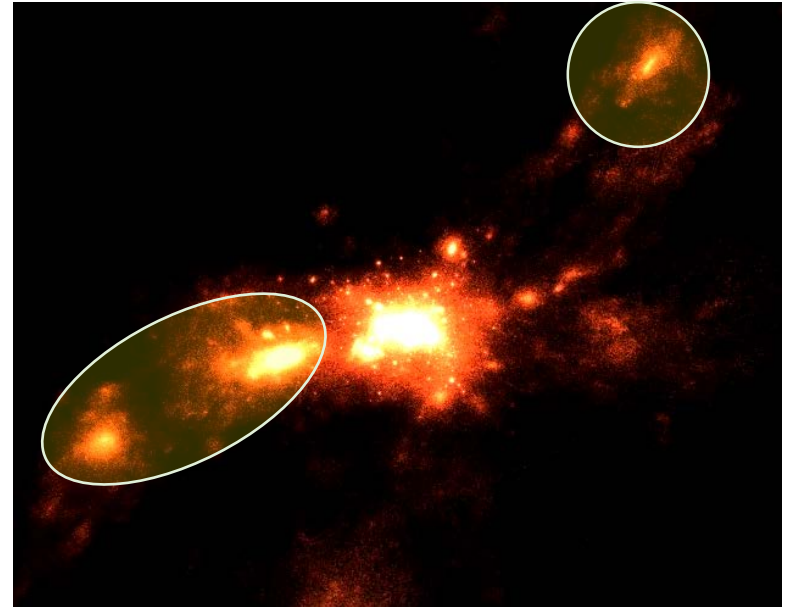
Earlier groups were more destructive!

Groups in different epoch



Early Groups

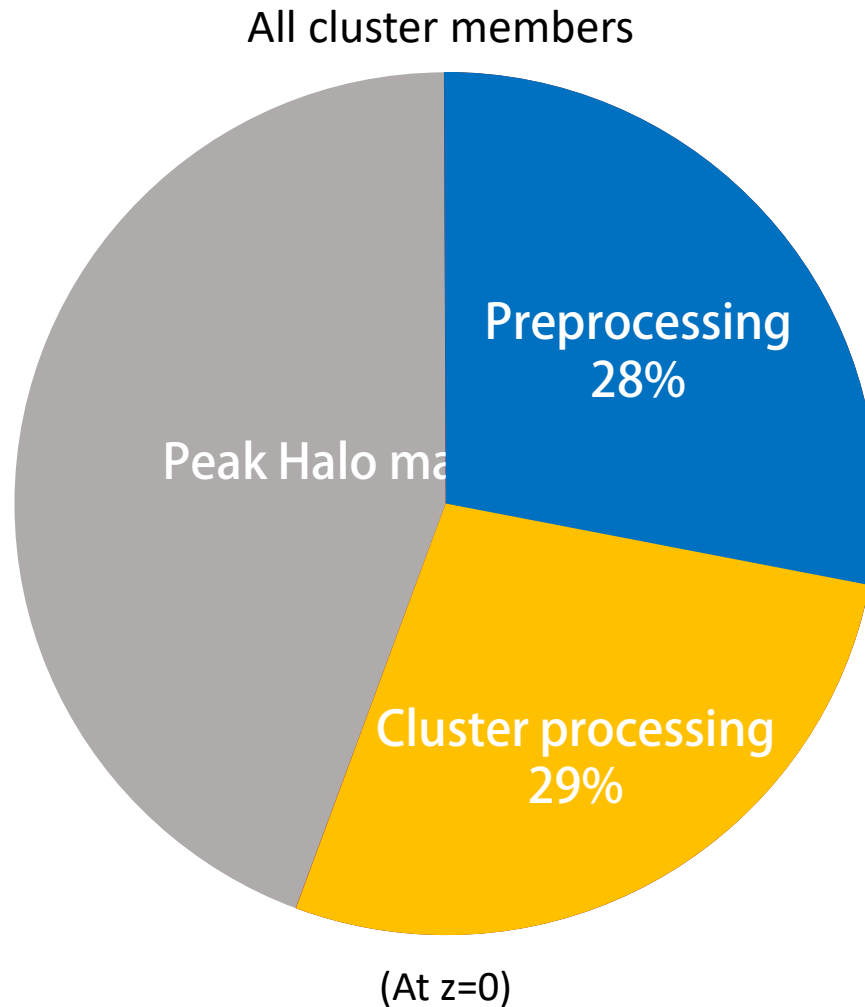
Violent mergers/accretions



Late Groups

Ordered accretion/growth

Mean mass loss of cluster members



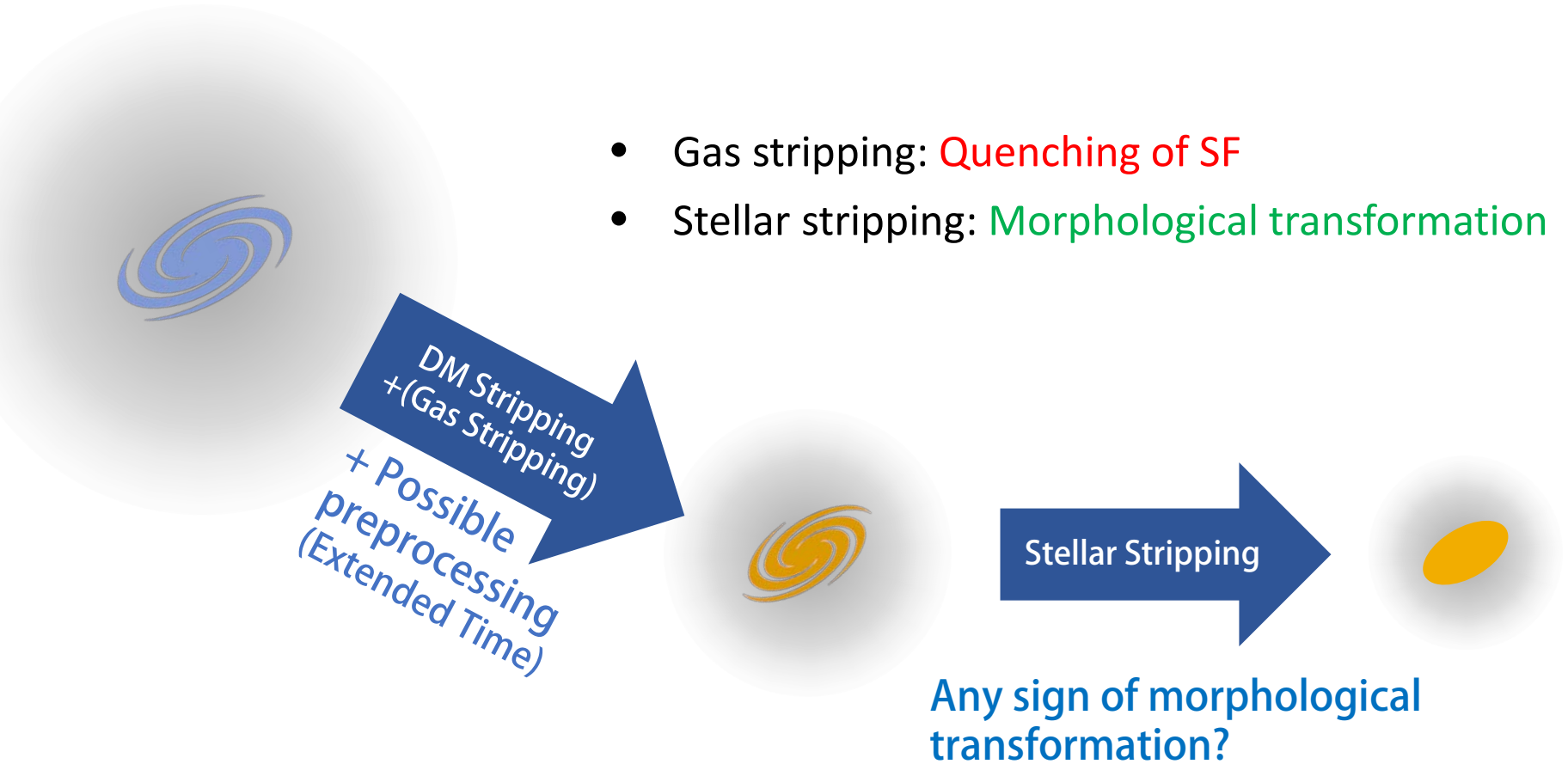
Summary

- Fraction of halos came from groups in the cluster $\sim 48\%$
- Higher preprocessed fraction = Recent rapid mass growth
- Rate of tidal stripping varies with:
 - host-satellite mass ratio
 - cosmic epoch.
- Tidal stripping in group is important as clusters!

Future Aspects

Transformation of galaxies

- Gas stripping: **Quenching of SF**
- Stellar stripping: **Morphological transformation**



Merci!